



## **Ray tracing calculations for the 3 pole wigglers**

**Manuel Sanchez del Rio**

## Introduction to Source modeling in SHADOW

Geometric

BM and Wiggler

Wiggler

Results for first 3-pole wiggler design

Results for new design

To do

Ray = array of 18 float (called columns – rows are number of rays):

( $\mathbf{x}$ ,  $\mathbf{v}$ ,  $\mathbf{E}_\sigma$ , lost\_flag,  $|\mathbf{k}|=2\pi/\lambda$ , index, OP,  $\phi_\sigma$ ,  $\phi_\pi$ ,  $\mathbf{E}_\pi$ )

3 3 3 1 1 1 1 1 1 3 = 18 => there is redundancy

We can easily define other variables (“compound columns”):

11:  $E$  [eV]

19:  $\lambda$  [Å]

20:  $r = (x^2 + y^2 + z^2)^{1/2}$

21: angle from Y axis

22:  $|E| = |E_\sigma + E_\pi|$

23:  $I = |E|^2$

24:  $I_\sigma = |E_\sigma|^2$

25:  $I_\pi = |E_\pi|^2$

26:  $k$

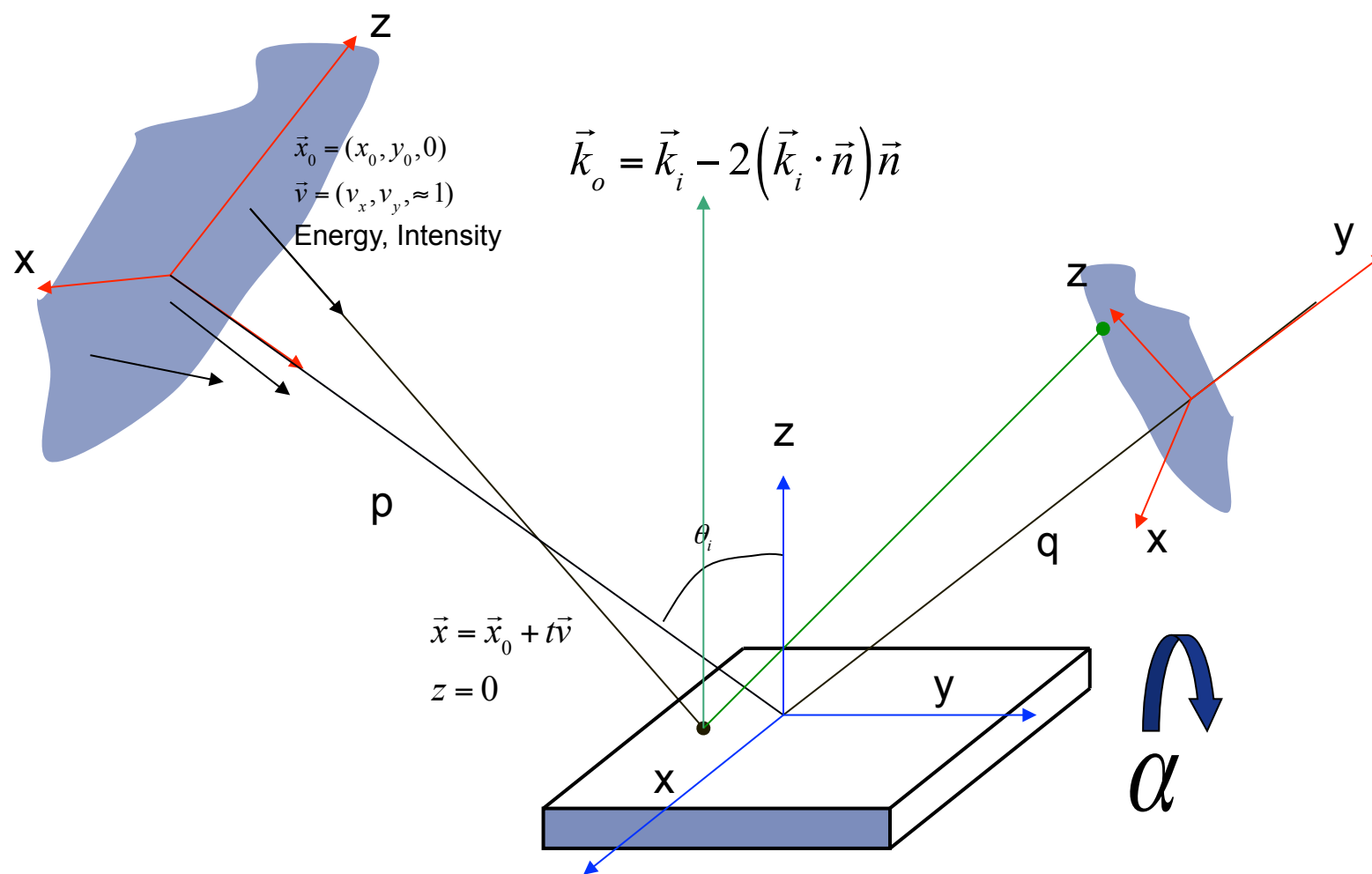
27-29: ( $k_x, k_y, k_z$ )

30-32 Stokes =  $\{ I_\sigma + I_\pi, I_\sigma - I_\pi, 2 |E_\sigma| |E_\pi| \cos(\phi_\sigma - \phi_\pi), 2 |E_\sigma| |E_\pi| \sin(\phi_\sigma - \phi_\pi) \}$

A single ray is always monochromatic

A beam is a collection of rays

# Trace (the beamline)



Fill with given distributions (flat, Gaussian, etc)  
 ( $\mathbf{x}$ ,  $\mathbf{v}$ ,  $\mathbf{E}_\sigma$ , lost\_flag,  $|\mathbf{k}|=2\pi/\lambda$ , index, OP,  $\phi_\sigma$ ,  $\phi_\pi$ ,  $\mathbf{E}_\pi$ )

E.g.,

point source  $\mathbf{x}=\{0,0,0\}$ ,

collimated source  $\mathbf{v}=\{0,0,1\}$

Note that

$|\mathbf{v}| = 1$ ,  $|\mathbf{E}_\sigma + \mathbf{E}_\pi|^2 = 1$  at the source

$\phi_\pi = \phi_\sigma + \text{phase difference}$

$\phi_\sigma$  is usually random (incoherent rays one to another)

In geometric sources,  $\mathbf{x}$ ,  $\mathbf{v}$  and  $\mathbf{E}$  are uncorrelated:

The sampling order is not important.

Once created the source, a variable can be resampled

$$(A_s)^2 + (A_p)^2 = 1$$

. SHADOW prompts the user for the phase difference in degrees between  $\phi_s$  and  $\phi_p$ . The Degree of polarization determines the relative amplitude of the  $s$  and  $p$  vectors according to the relation :

$$\text{Degree of polarization} = \cos(\theta)/(\cos(\theta) + \sin(\theta))$$

where  $\theta$  is the angle of polarization from the  $A_s$  axis.

Some examples of polarization states:

- Linearly polarized:

Phase difference = 0

Degree of polarization = whatever the above expression is when the desired value of  $\theta$  is plugged in.

for  $\theta = 45$  deg,  $DOP = 0.50$

for  $\theta = 30$  deg,  $DOP = 0.63$

- Circularly polarized:

Phase diff = +90 (for Right), -90 (for Left)

Degree of Polarization = 0.5

# Monte Carlo (source model)

INVERSION

REJECTION

THE INSTITUTE FOR ADVANCED STUDY  
SCHOOL OF MATHEMATICS  
PRINCETON, NEW JERSEY

May 21, 1947

Mr. Stan Ulam  
Post Office Box 1663  
Santa Fe  
New Mexico

Dear Stan,

Thanks for your letter of the 19th. I need not tell you that Klari and I are looking forward to the trip and visit at Los Alamos this Summer. I have already received the necessary papers from Carson Mark. I filled out and returned mine yesterday; Klari's will follow today.

I am very glad that preparations for the random numbers work are to begin soon. In this connection, I would like to mention this: Assume that you have several random number distributions, each equidistributed in  $0, 1$ :  $(x^i), (y^i), (z^i), \dots$ . Assume that you want one with the distribution function (density)  $f(x) dx = f(x^i)$ . One way to form it is to form the cumulative distribution function:  $g(x) = \int_0^x f(t) dt$  to invert it  $g(x) = \xi \Rightarrow x = g^{-1}(\xi)$ , and to form  $\xi^i = g^{-1}(x^i)$  with this  $g^{-1}$ , or some approximant polynomial. This is, as I see, the method that you have in mind.

An alternative, which works if  $f$  and all values of  $f(x)$  lie in  $0, 1$ , is this: Scan pairs  $x^i, y^i$  and use or reject  $x^i, y^i$  according to whether  $y^i \leq f(x^i)$  or not. In the first case, put  $\xi^i = x^i$  in the second case form no  $\xi^i$  at that step.

The second method may occasionally be better than the first one. In some cases combinations of both may be best; e.g., form random pairs  $\xi = \sin x, \eta = \cos x$  with  $x$  equidistributed between  $0^\circ$  and  $360^\circ$ . The obvious way consists of using the sin - cos - tables (with interpolation). This is clearly closely related to the first method. This is an alternative procedure:

Put  $\xi = \frac{2t}{1+t^2}, \eta = \frac{1-t^2}{1+t^2}, t = \tan y$ , with  $y$  (which is  $\frac{x}{2}$ ) equidistributed between  $0^\circ$  and  $180^\circ$ . Restrict  $y$  to  $0^\circ$  to  $45^\circ$ . Then the  $\xi, \eta$  will have to be replaced randomly by  $\eta, \xi$  and again by  $\pm \xi, \pm \eta$ . This can be done by using random digits  $0, \dots, 7$ . It is also feasible with

- 1) Sample Energies from BM spectrum (pre-calculated or exact)
- 2) Sample position (uniformly distributed over the trajectory arc)
- 3) Sample direction
  - Tangent to trajectory in H
  - Correlated to photon energy: Sample angle following, get DOP from  $|E_s|/|E_p|$  ratio

Add emittance effects:

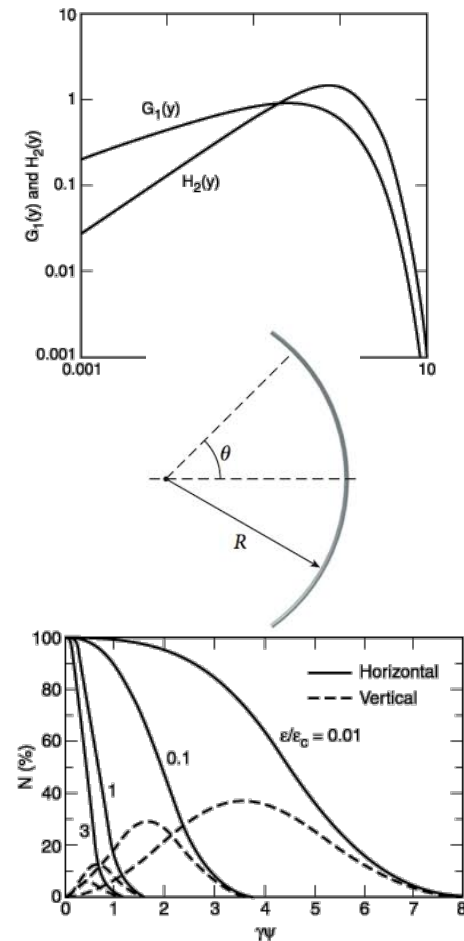
- 1) Ask  $\sigma, \sigma'$  and waist position (for H and V)
- 2) Compute e<sup>-</sup> beam size distribution at the ray starting position
- 3) Sample a random value using this bivariate Gaussian distribution (for H and V)
- 4) Add it to the ray position  $\mathbf{x}$  and direction  $\mathbf{v}$

Good:

Exact model for emission, good geometry characteristics (depth, curvature),

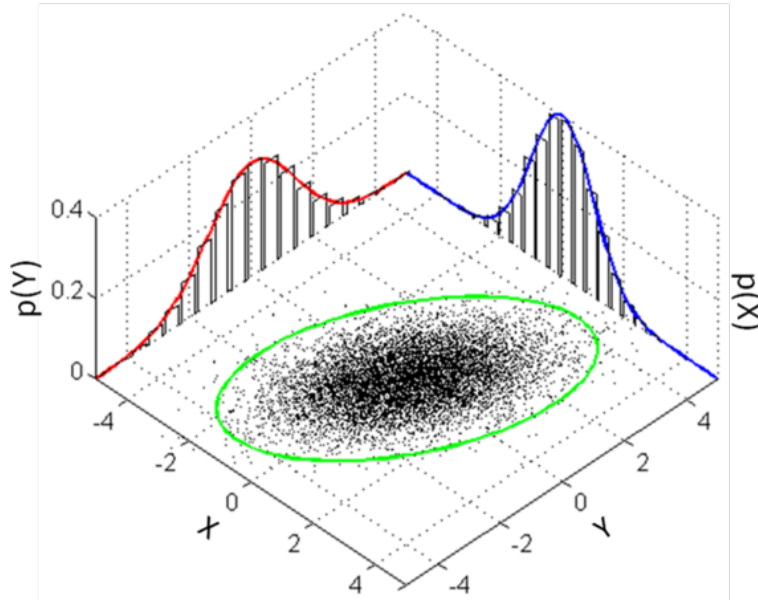
Limitations:

- i) No edge radiation
- ii) the e<sup>-</sup> beam size at a given position  $s$  is obtained from the waist position supposing we are in free space (no magnetic field)  
=> small effect in general => to be checked and upgraded!
- iii) No energy spread



## COMPUTE E- BEAM SIZES

At **s** (any point of the trajectory):



### Bivariate Normal Distribution

$$\Sigma = \begin{pmatrix} \langle x^2 \rangle & \langle xx' \rangle \\ \langle xx' \rangle & \langle x'^2 \rangle \end{pmatrix} = \begin{pmatrix} \beta_x \varepsilon_x & -\alpha_x \varepsilon_x \\ -\alpha_x \varepsilon_x & \gamma_x \varepsilon_x \end{pmatrix} + \eta^2 \sigma_\delta^2 I_{2 \times 2}$$

### Evolution in empty space

$$\begin{aligned} \langle x^2 \rangle_y &= \langle x^2 \rangle + 2 \langle xx' \rangle y + \langle x'^2 \rangle y^2 \\ \langle xx' \rangle_y &= \langle xx' \rangle + \langle x'^2 \rangle y \\ \langle x'^2 \rangle_y &= \langle x'^2 \rangle \end{aligned}$$

With  $\varepsilon$  the emittance (constant), and Twiss parameters:

$$\sigma_x = \sqrt{\langle x^2 \rangle} = \sqrt{\beta_x \varepsilon_x}; \quad \sigma_{x'} = \sqrt{\langle x'^2 \rangle} = \sqrt{\gamma_x \varepsilon_x}; \quad \sigma_x \sigma_{x'} = \varepsilon_x \sqrt{1 + \alpha_x^2} \quad \alpha = -\frac{1}{2} \frac{d\beta}{ds}; \quad \gamma = \frac{1 + \alpha^2}{\beta}$$

At **waist** (zero correlation,  $\rho = \alpha = 0$ ,  $\beta$  is minimum):

$$\sigma_x = \sqrt{\langle x^2 \rangle} = \sqrt{\beta_x \varepsilon_x}; \quad \sigma_{x'} = \sqrt{\langle x'^2 \rangle} \Big|_w = \sqrt{\frac{\varepsilon_x}{\beta_x}}; \quad \boxed{\sigma_x \sigma_{x'} = \varepsilon_x}$$





# WIGGLER MODEL IN SHADOW: SINGLE ELECTRON

Calculate

Spectrum (full emission)

Trajectory

Velocities (directions or angles)

Curvature ( $1/R$ )

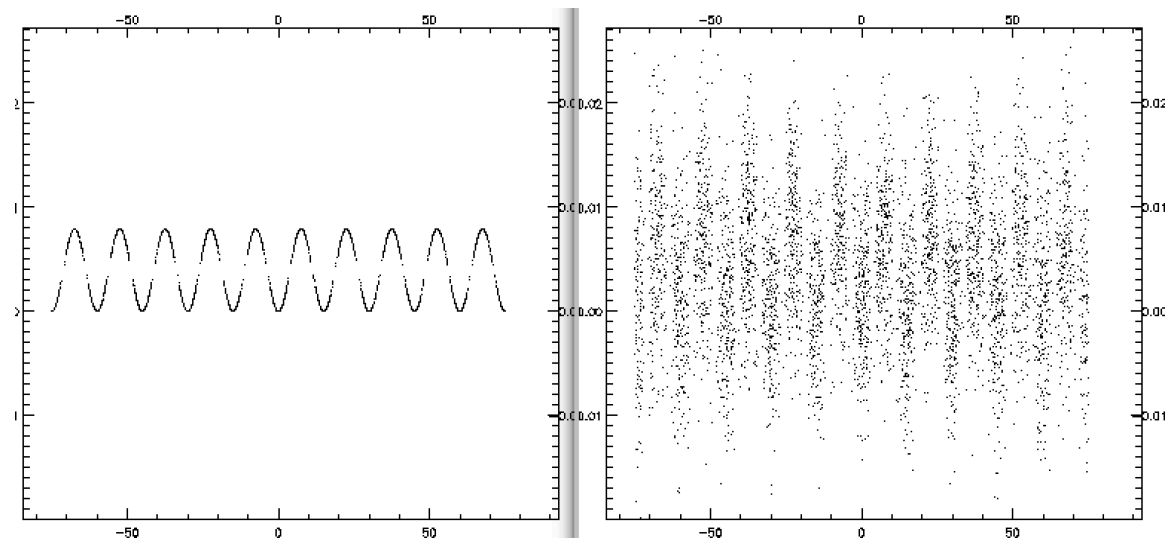
Sample:

photon energy

position in the trajectory (not uniform, sampling from curvature)

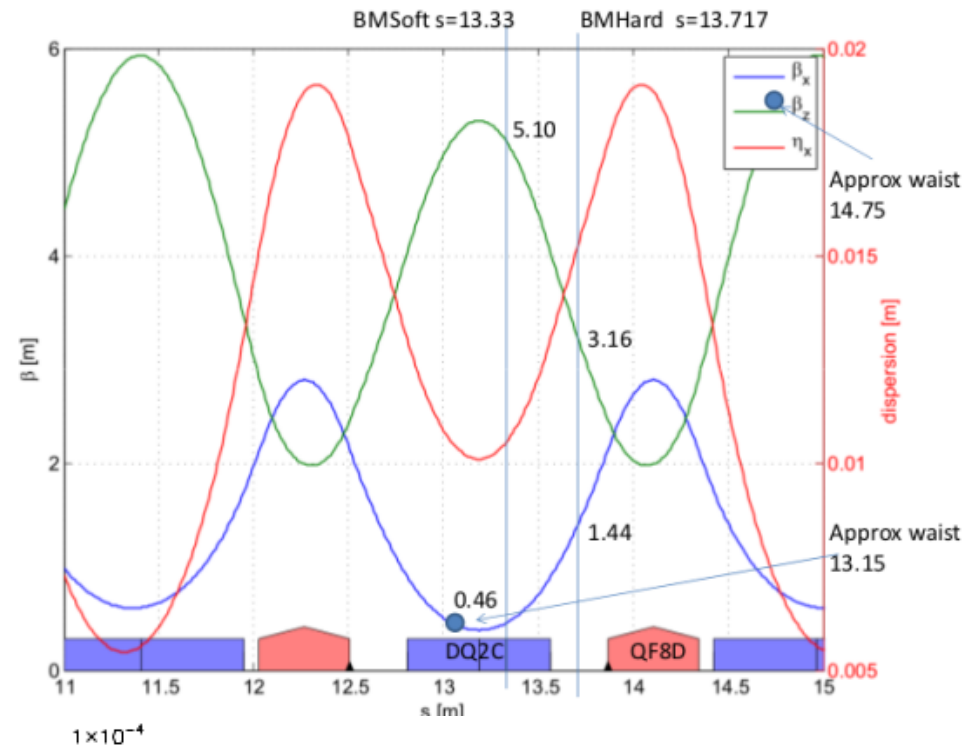
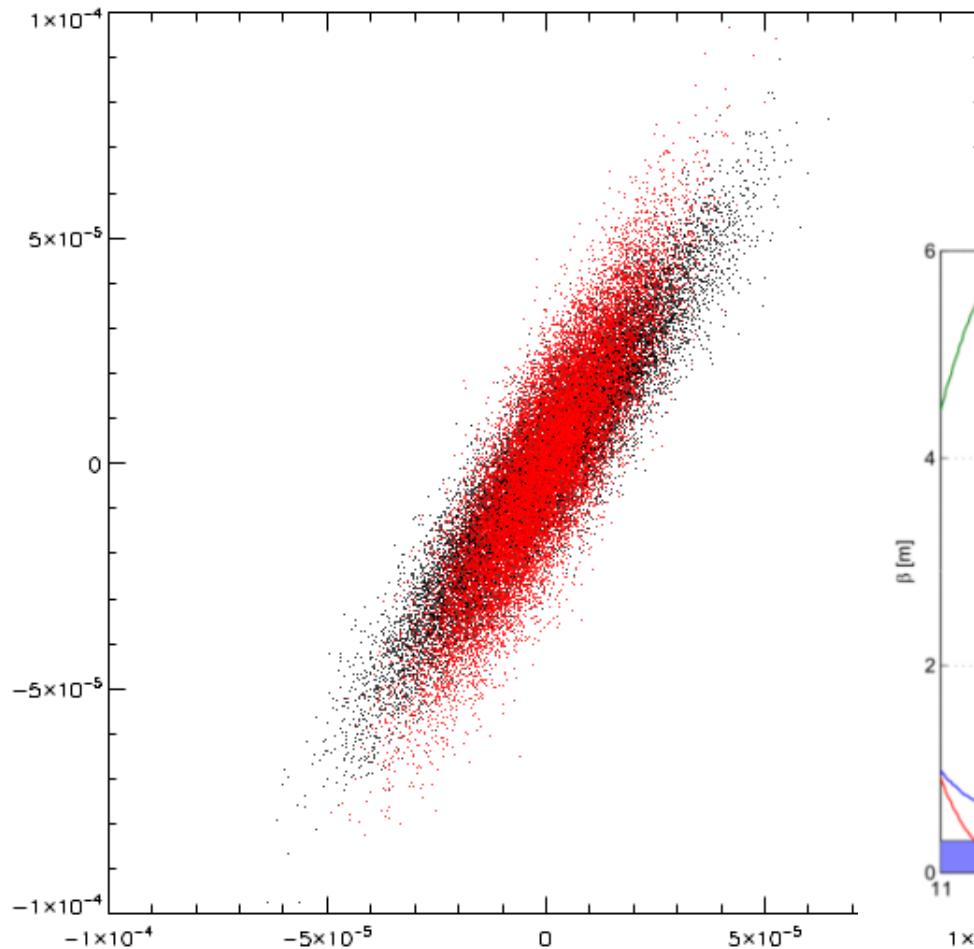
angle: H tangent to trajectory

V: Sampled from BM distributions for the local magnetic radius



# WIGGLER MODEL IN SHADOW: ELECTRON PARAMETERS

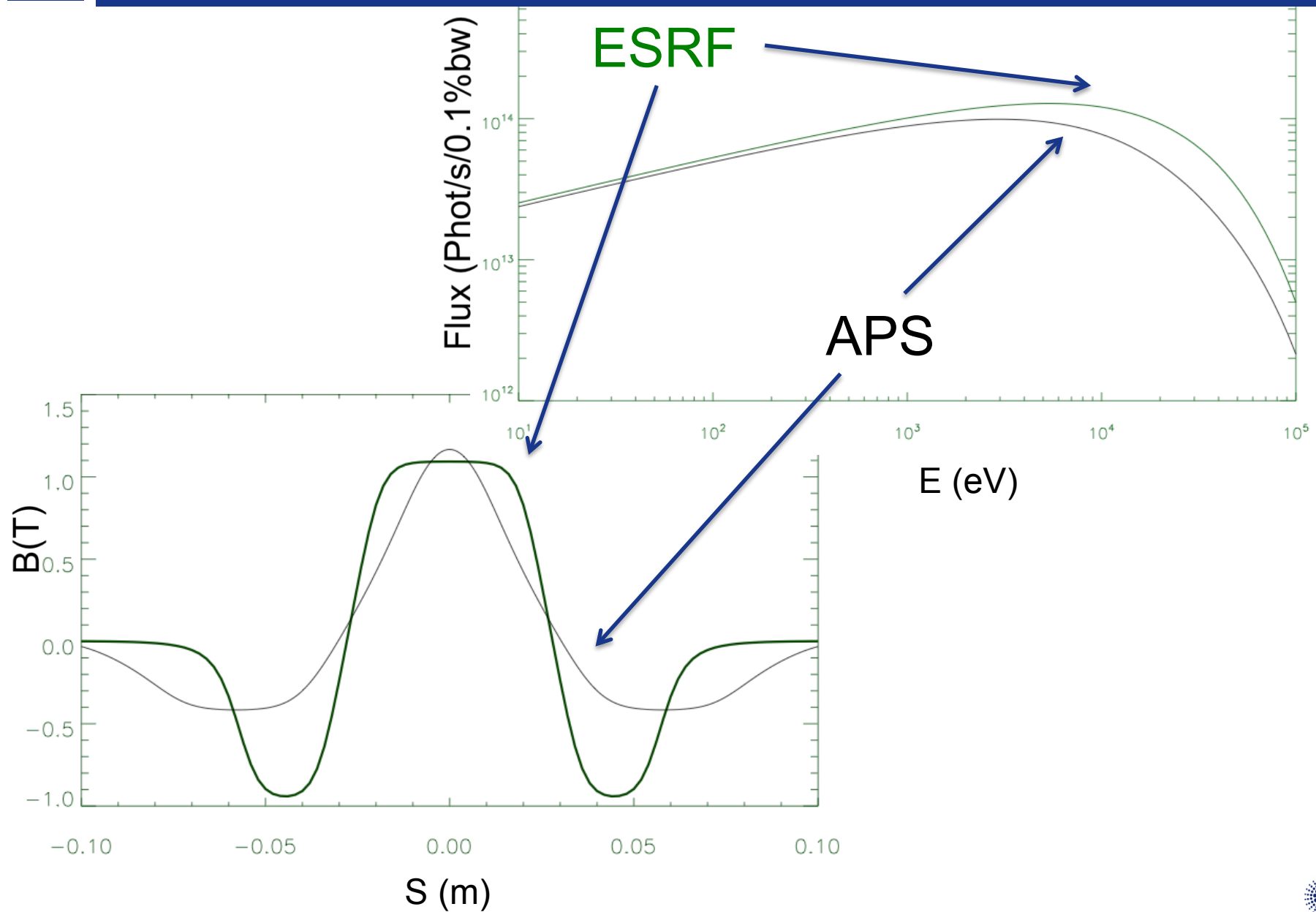
3PW at the ESRF new lattice



Monte Carlo sampling of the electron phase space  $x'[\text{rad}]$  vs.  $x[\text{m}]$  at wiggler edges:  $y = -0.1$  (red) and  $y = 0.1$  (black).

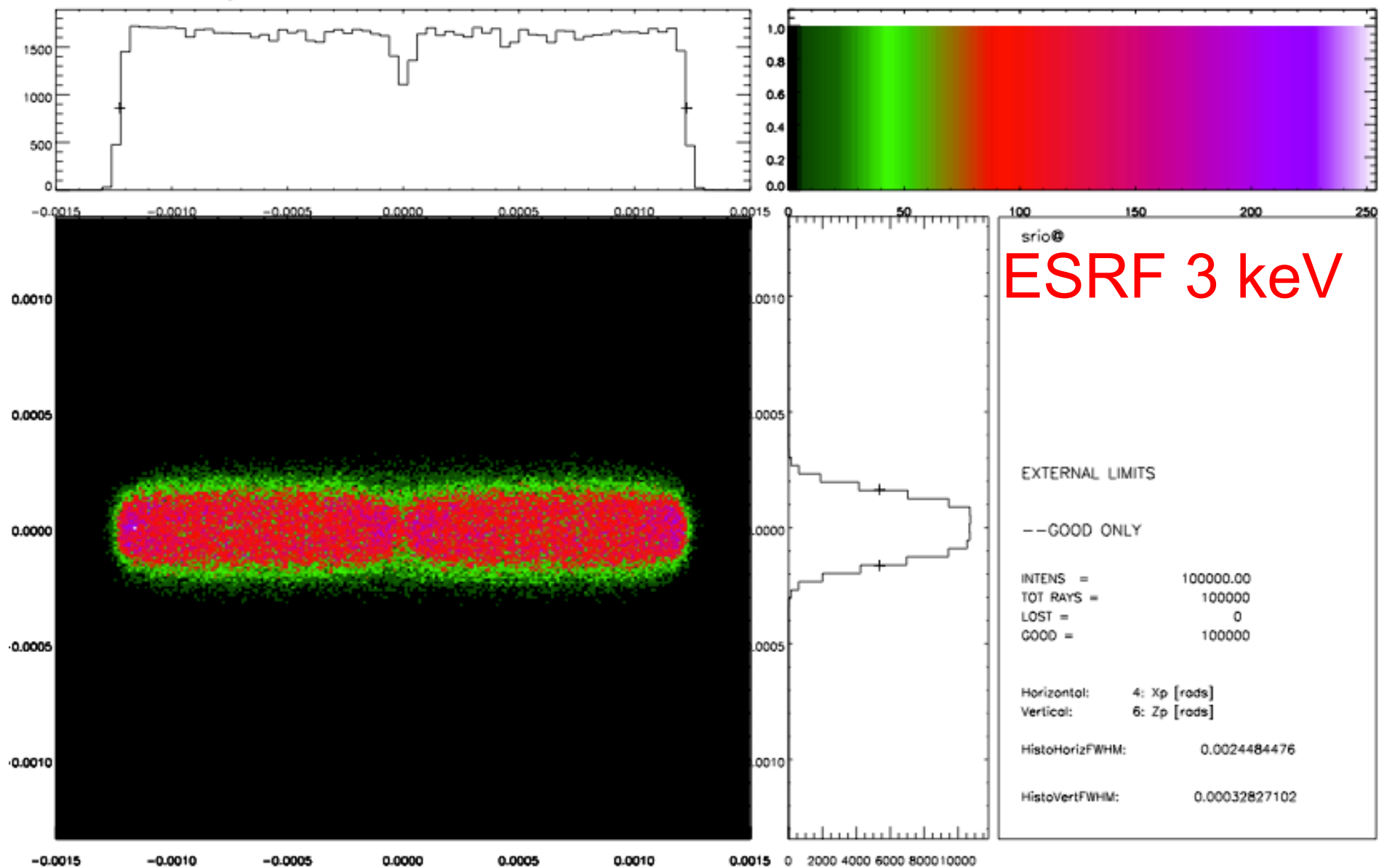
Implementation in SHADOW: <http://forge.epn-campus.eu/documents/784>

## 3PW COMPARISON



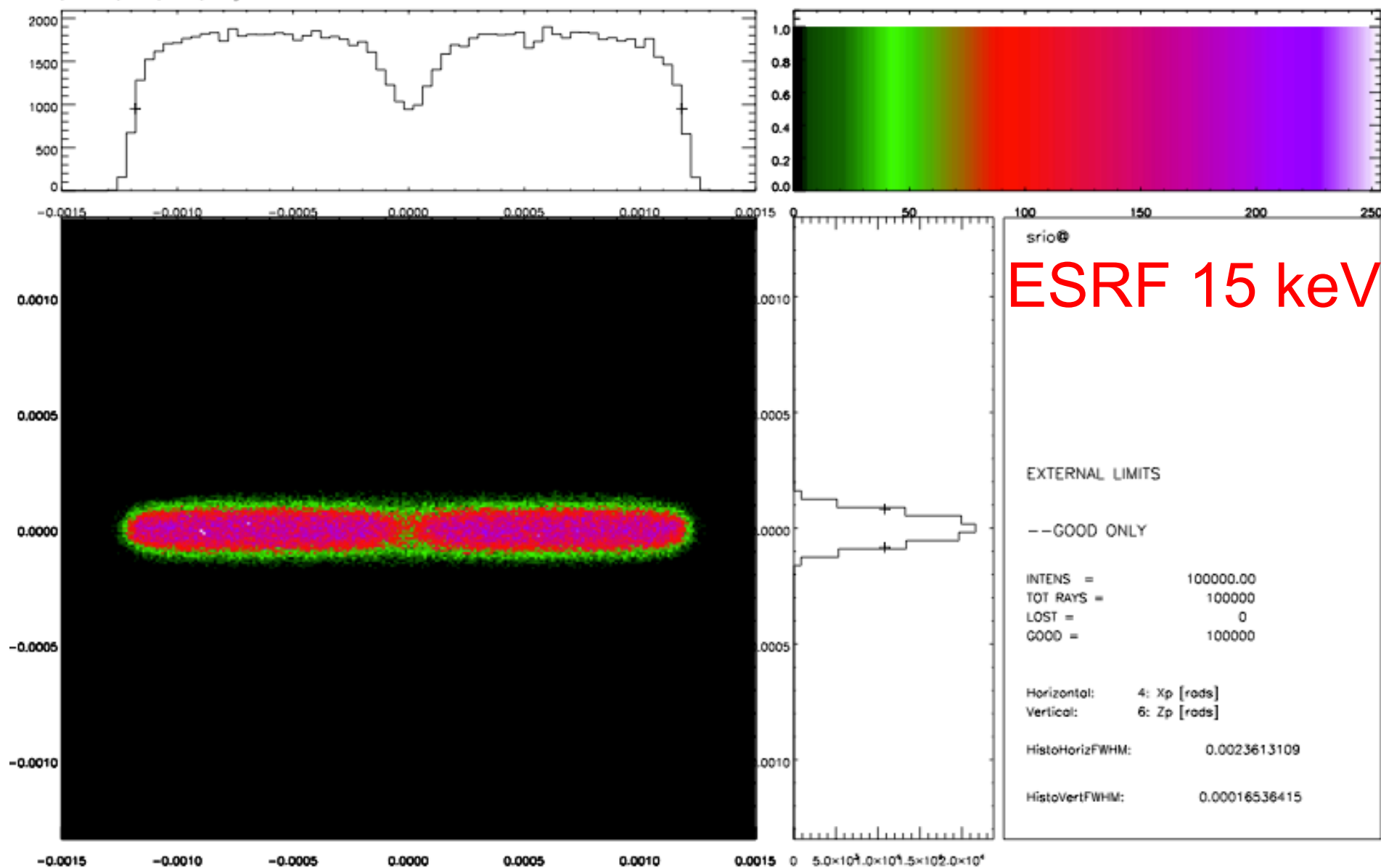
command: plotxy,begin,dot,4,6,XRANGE=[-0.00150000,0.00150000],NOLOST=1,NBINS=75,CALFWHM=1,CONTOUR=5,NLEVELS=6

/Users/srio/TMP/begin.dot Ven 21 nov 2014 17:05:16 CET



command: plotxy,begin,dot,4,6,XRANGE=[-0.00150000,0.00150000],NOLOST=1,NBINS=75,CALFWHM=1,CONTOUR=5,NLEVELS=6

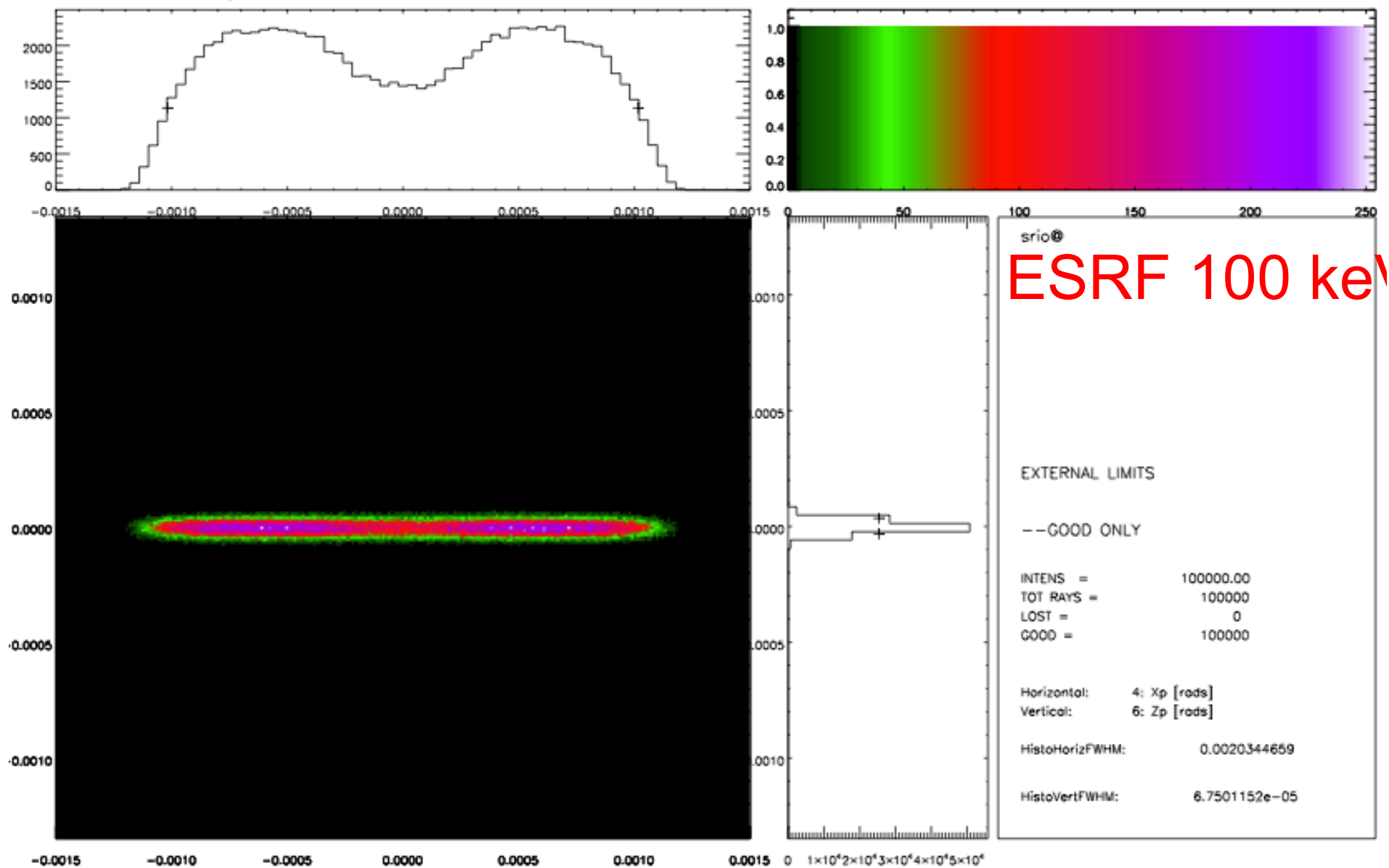
/Users/srio/TMP/begin.dot Ven 21 nov 2014 17:07:40 CET



ESRF 15 keV

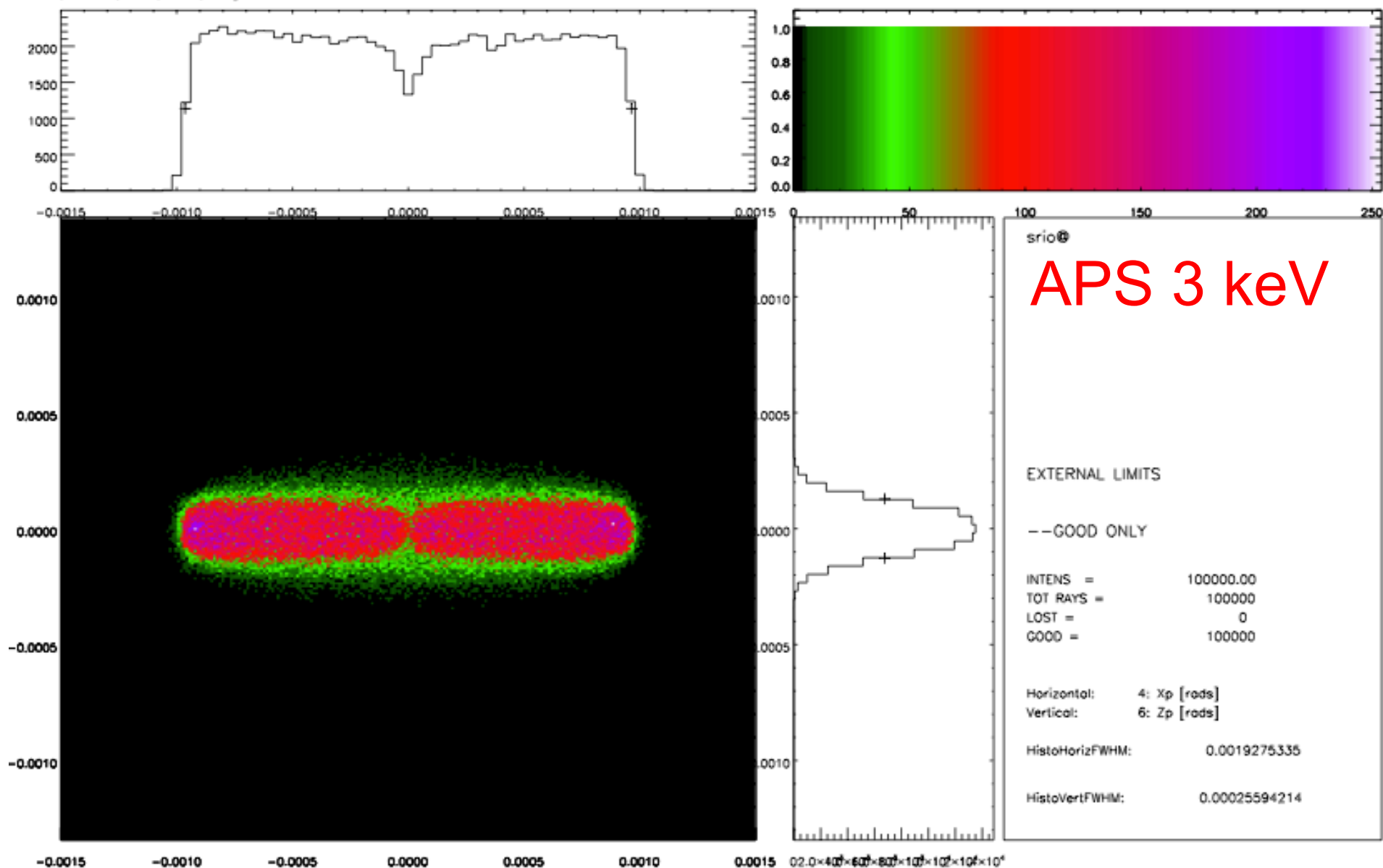
```
command: plotxy,begin.dat,4,6,XRANGE=[-0.00150000,0.00150000],NOLOST=1,NBINS=75,CALFWHM=1,CONTOUR=5,NLEVELS=6
```

/Users/srio/TMP/begin.dat Ven 21 nov 2014 17:09:13 CET



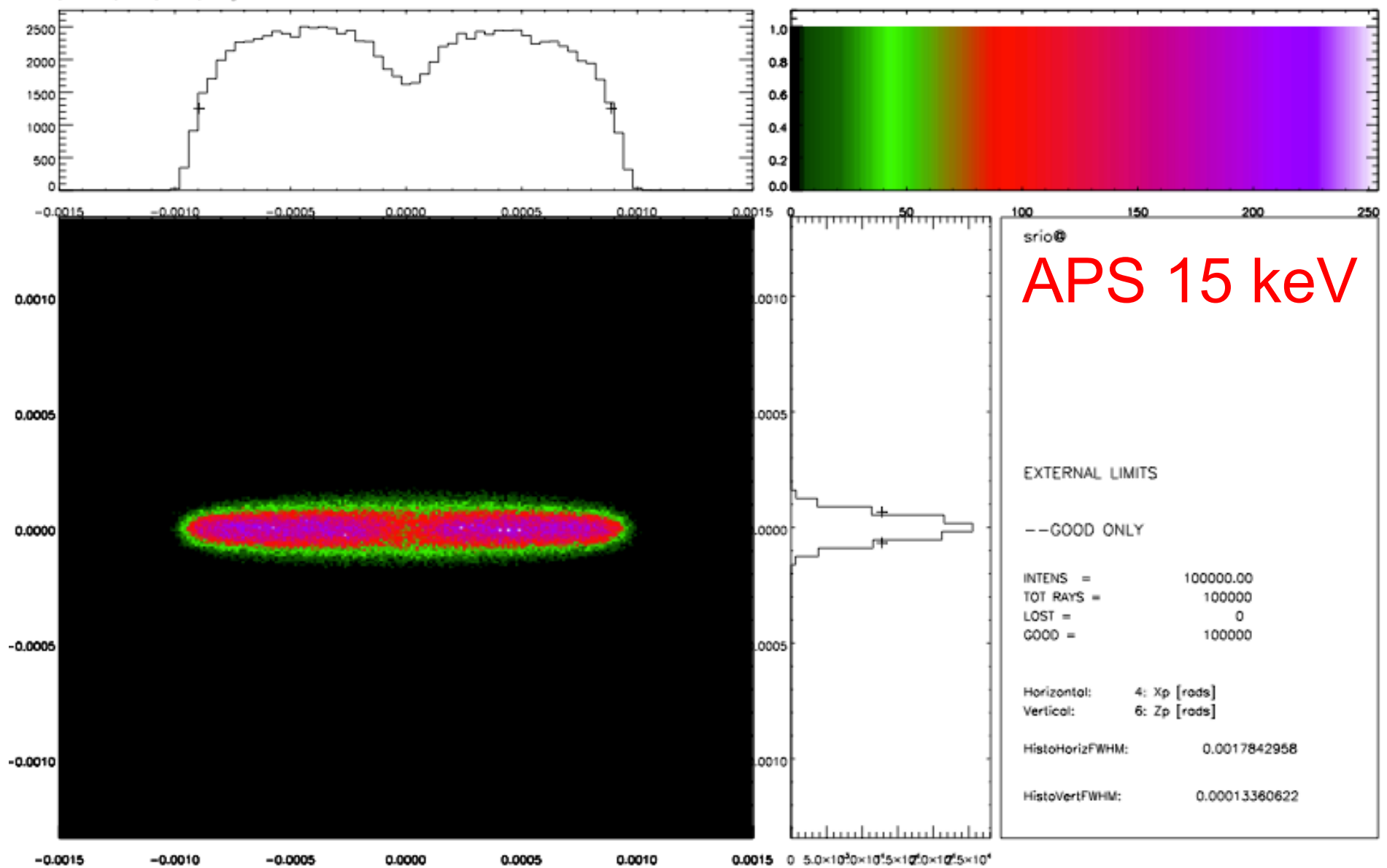
command: plotxy,begin.dat,4,6,XRANGE=[-0.00150000,0.00150000],NOLOST=1,NBINS=75,CALFWHM=1,CONTOUR=5,NLEVELS=6

/Users/srio/TMP/begin.dat Ven 21 nov 2014 17:28:32 CET



command: plotxy,begin,dot,4,6,XRANGE=[-0.00150000,0.00150000],NOLOST=1,NBINS=75,CALFWHM=1,CONTOUR=5,NLEVELS=6

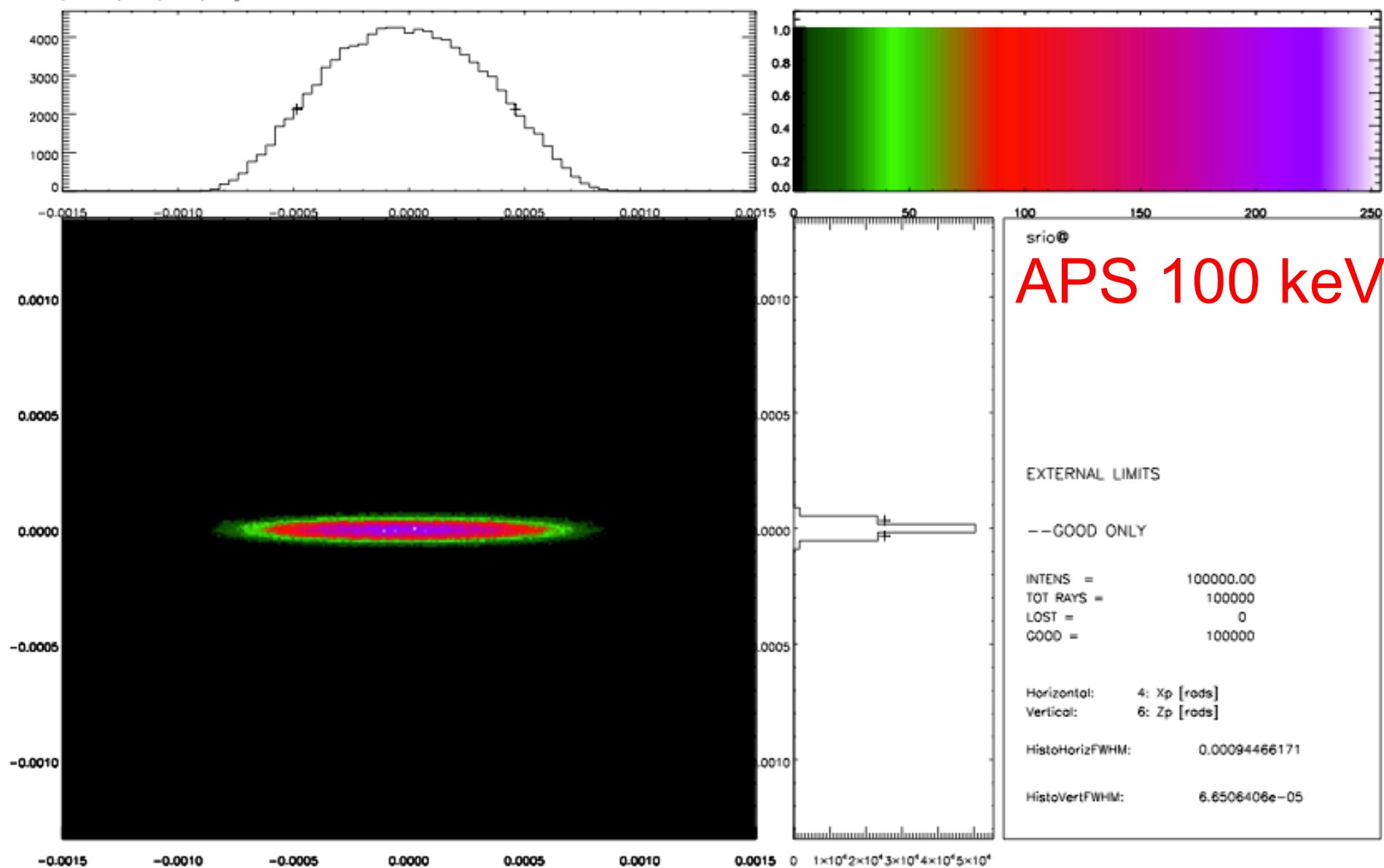
/Users/srio/TMP/begin.dat Ven 21 nov 2014 17:27:02 CET





command: plotxy,begin.dat,4,6,XRANGE=[-0.00150000,0.00150000],NOLOST=1,NBINS=75,CALFWHM=1,CONTOUR=5,NLEVELS=6

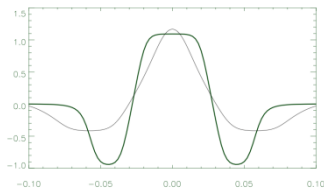
/Users/srio/TMP/begin.dat Ven 21 nov 2014 17:25:58 CET



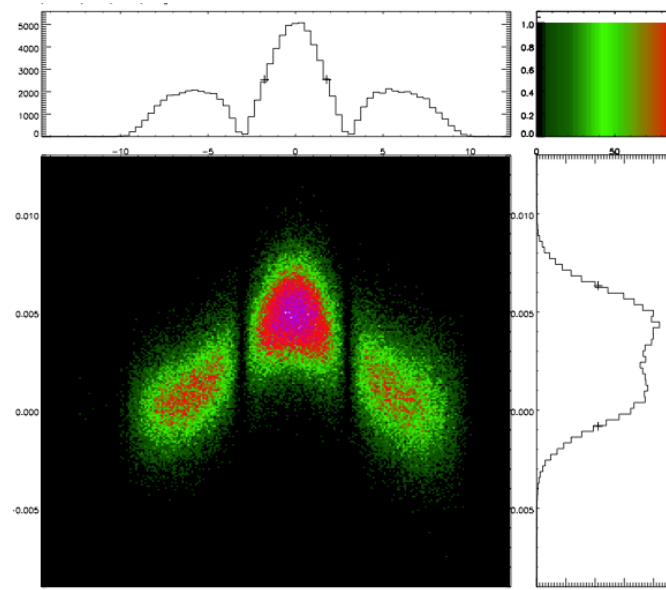
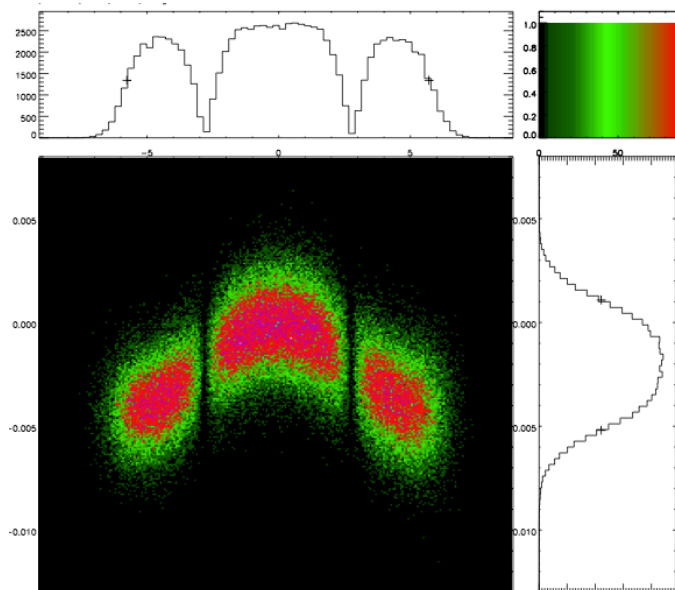
# Emission – Top view

ESRF

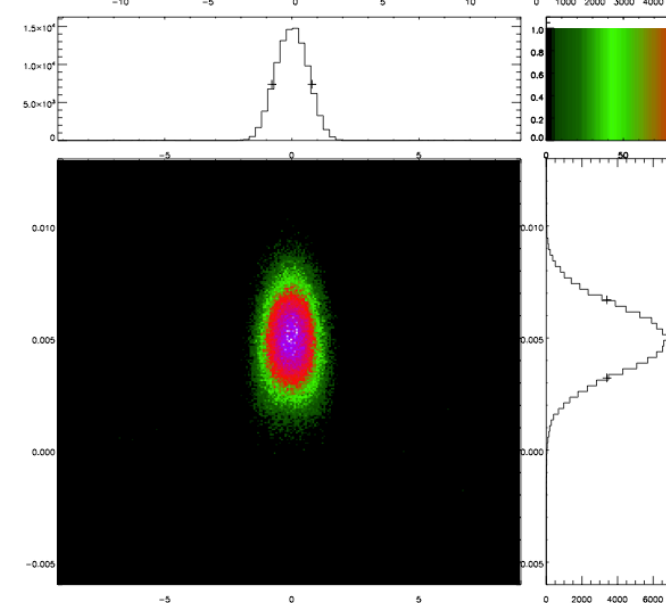
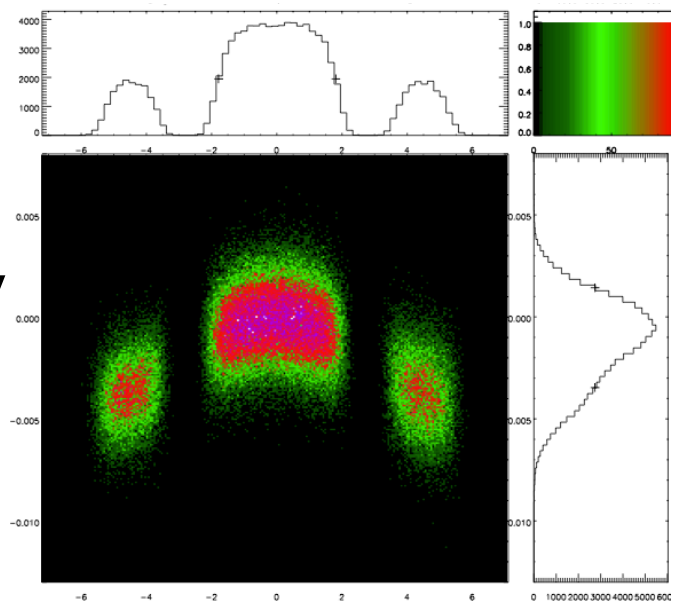
APS



3 keV



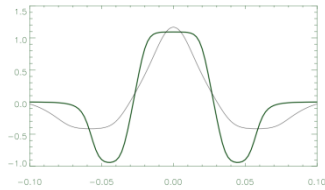
100 keV



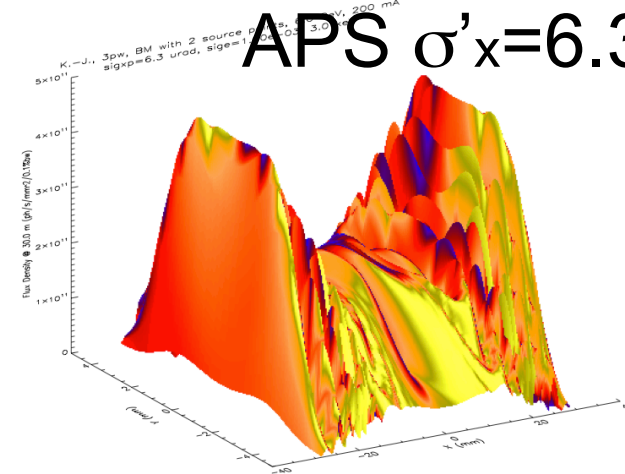
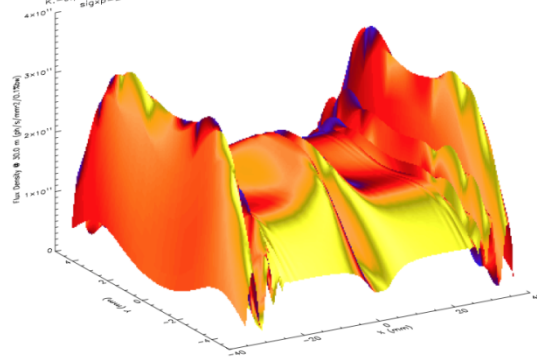
# Beam profile @ 30m

ESRF  $\sigma'_x = 23 \mu\text{rad}$

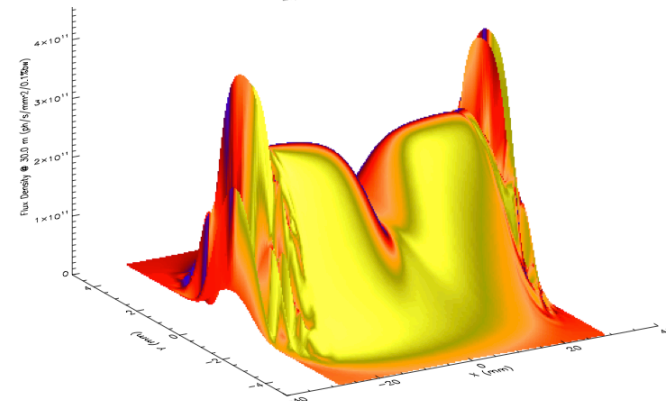
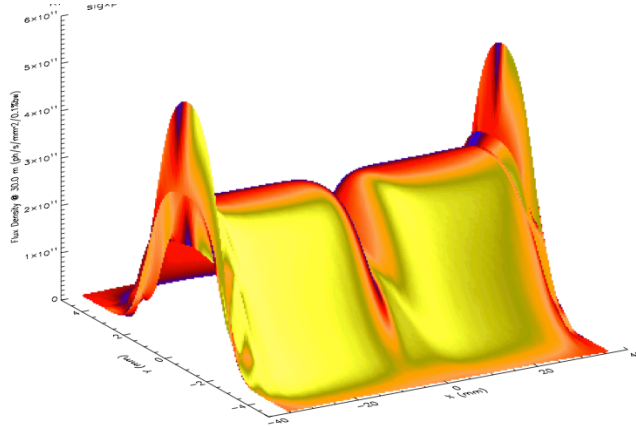
APS  $\sigma'_x = 6.3 \mu\text{rad}$



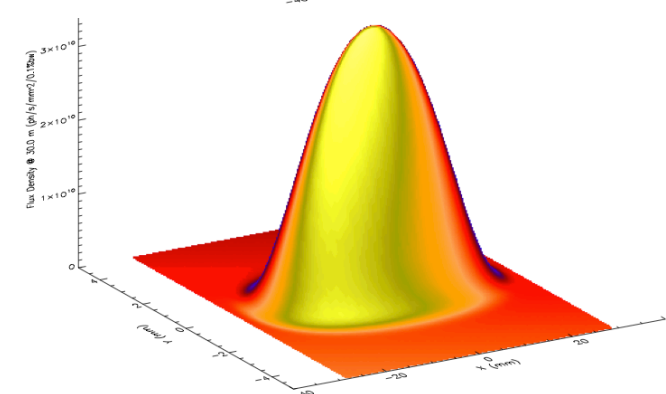
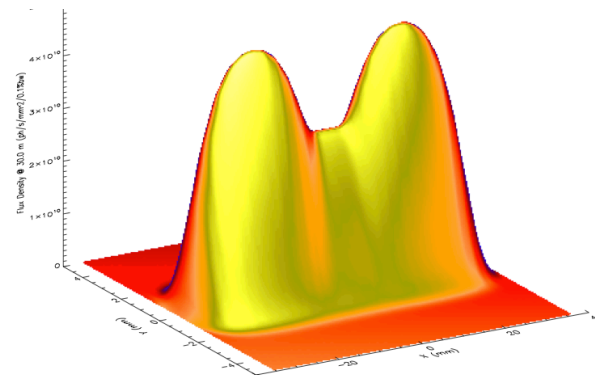
3 keV



15 keV



100 keV



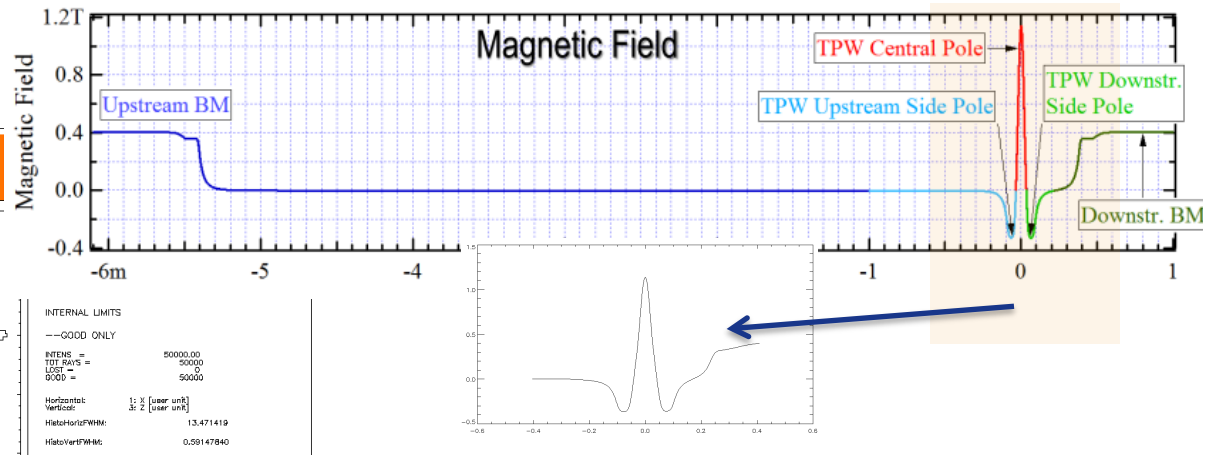
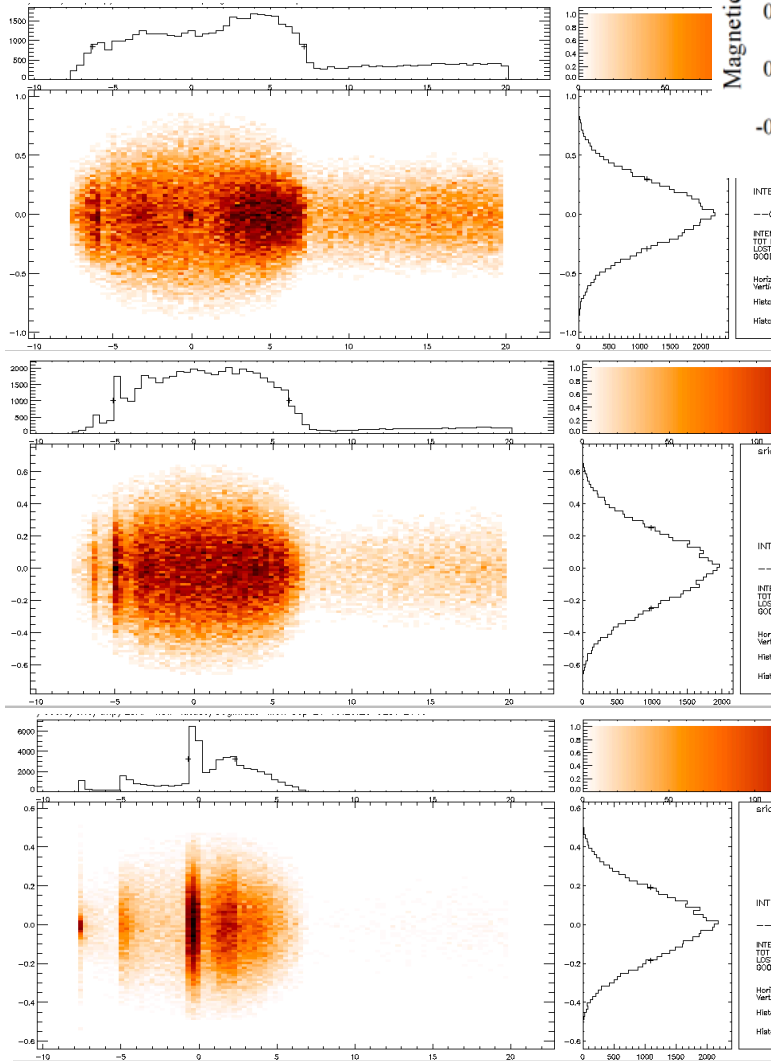
Kwang-Je Kim and Roger Dejus

Sharpness of Interference Pattern of APSU 3-Pole Wiggler - MD-TN-2014-003

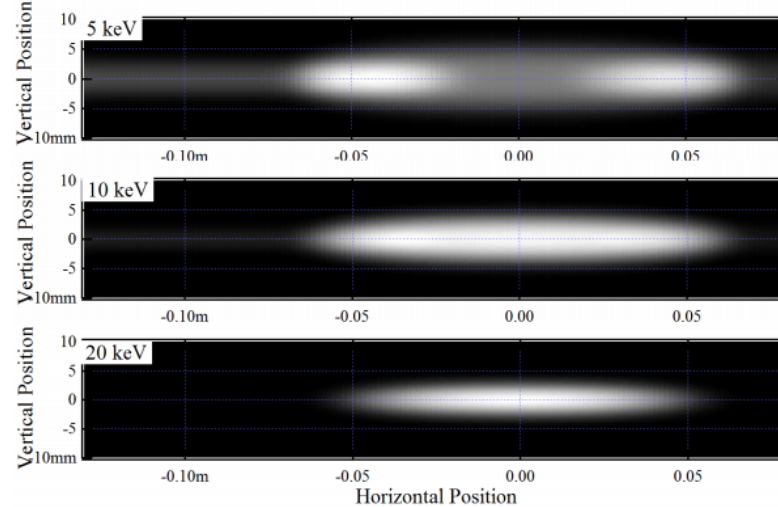
The European Synchrotron



# NSLS-II 3PW SHADOW-NO EMITTANCE



## 3PW and BM Intensity Distributions (Hard X-rays)



- Intensity distributions at different photon energy 30 m from 3PW show a broad distribution from soft poles in 3PW and a more focused distribution from adjacent BMs
- Effect of such non-ideal intensity distribution or microfocusing is being studied by a working group, and updates will be provided

•Calculation with DEJUS?

## 3PW – OVERLAPPING WITH BM

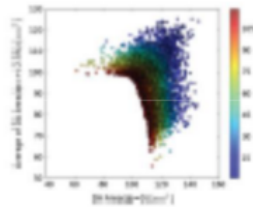
- There is an overlap with the BM radiation of the downstream BM
- Adding BM fields in SHADOW. Problems:
  - The axes for SHADOW are in the direction of the entrance electron, so the optical axis is not correctly placed
  - Emittance values not correctly set because is not empty space but there is magnetic fields (BMs, QP) [not dramatic...]
  - Still the problem of interpolation errors...
- Shortcuts:
  - Shift (to the e- starting position) and rotate (center of the wiggler at zero angle) the source

- PRESENT WAY:
  - Look for values published (reports, web, ...) and copy them
  - Talk to accelerator people
  - Values change... Tabulations are incomplete, and do not include all “interesting points”
- NEW WAY
  - Ring parameters from accelerator codes (internal formats...)
  - Run codes to obtain a tabulation of the Twiss parameters, moments, etc. at the desired  $[s_1, s_2]$  interval
  - Use this as input to sample rays in SHADOW, etc.

# CALCULATE/IMPORT ELECTRON BEAM CHARACTERISTICS

ALS

## Development of accelerator simulation tools



Symplectic Tracking based methods

DA, MA separated

DA, MA together



from L. Nadolski,  
ICFA LowEring,  
Oxford 7/13

Direct tracking based optimization

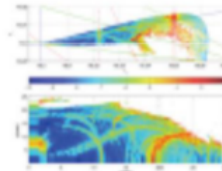
GLASS

Analytical based method

Genetic Algorithm  
MOGA

Lie Algebra/Differential Algebra

Frequency Maps  
FMA  
Diffusion factor



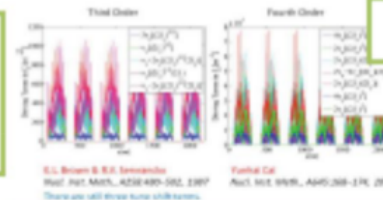
Nonlinear  
"LOCO"

Resonance Driving Terms  
RDT minimization

Amplitude Tunes shift  
minimization

Resonance identification

Canceling  
Sextupole  
Resonances



Phase advances

Interleaved  
sextupoles

Robustness to magnetic, alignment errors

Robustness ID configurations

Tracking codes: PTC MADXTRAY AT LIGO OPA ELEGANT

Advanced Light Source

D. Robin SLAC-DLSR Workshop



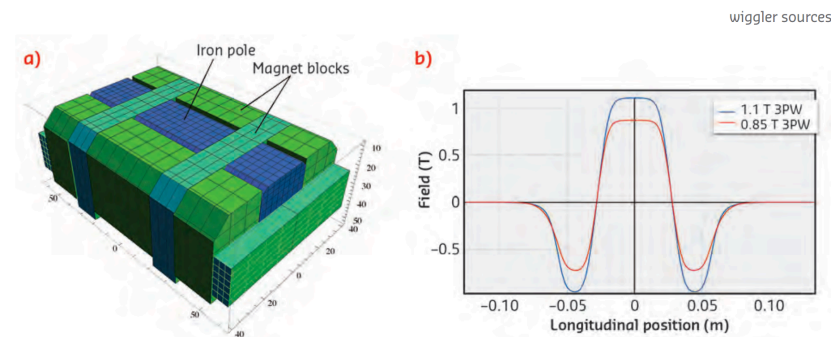


## EXAMPLE AT TO SHADOW

I want in a file with:

s B <xx> <xx'> <x'x'> <yy> <yy'> <y'y'>

Magnetic field:



AT provided these *plots* :

`load("S28CINJ.mat")`

`atplot(ARCA)`

Boaz helped to make a

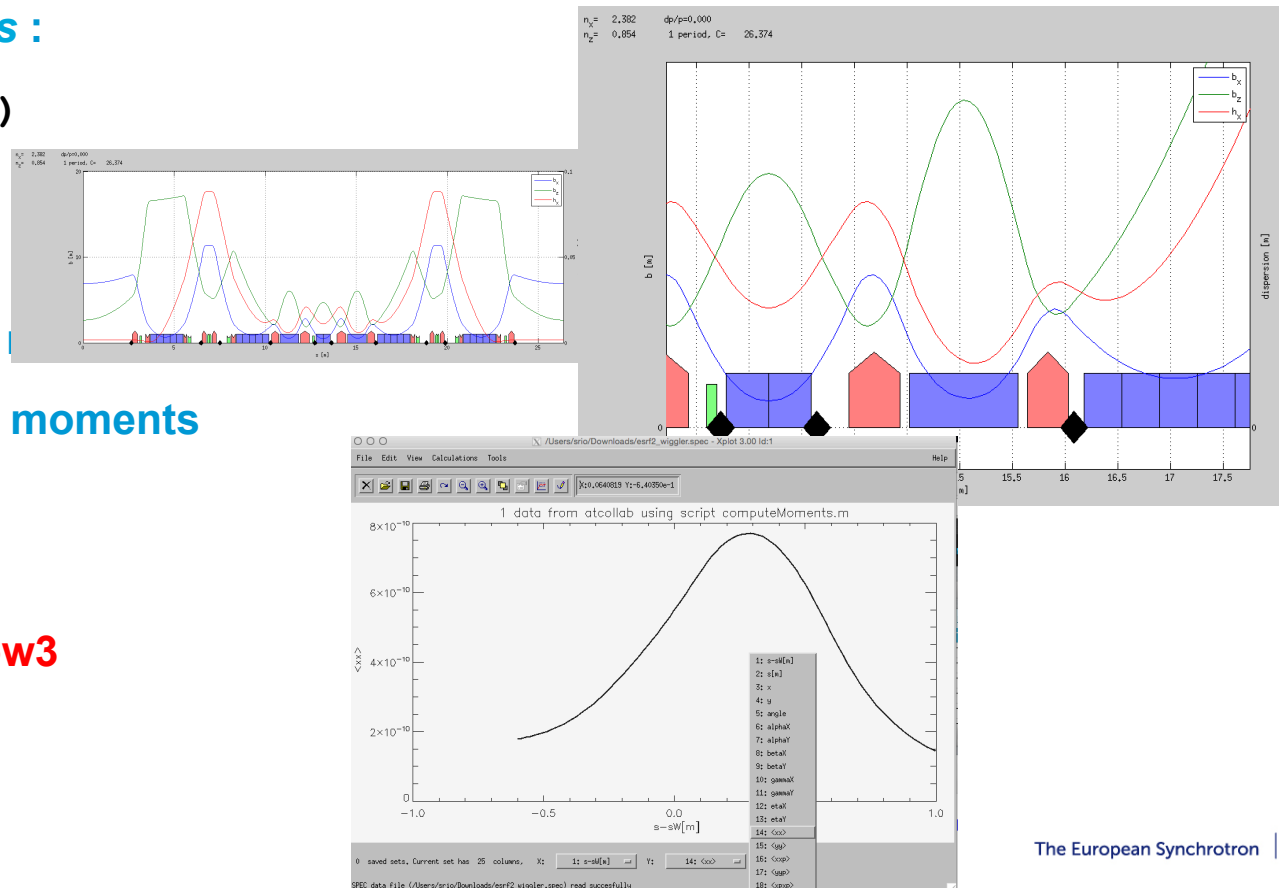
to drop Twiss pars and moments

**TODO:**

**Match them!!**

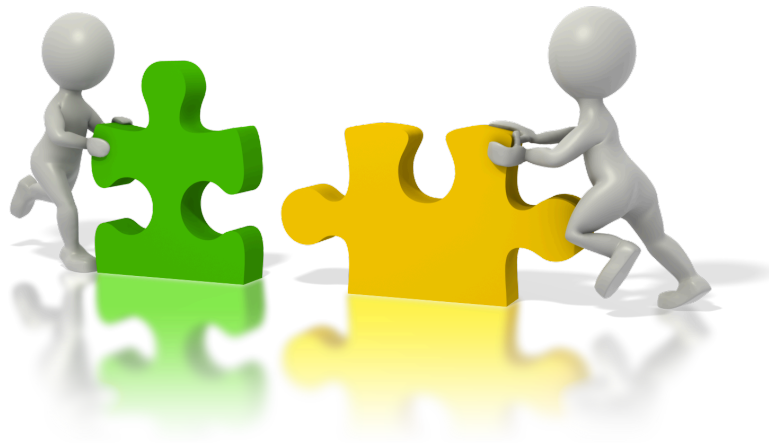
**Modify shadow3**

**Automatize?**





THE END



Thanks!